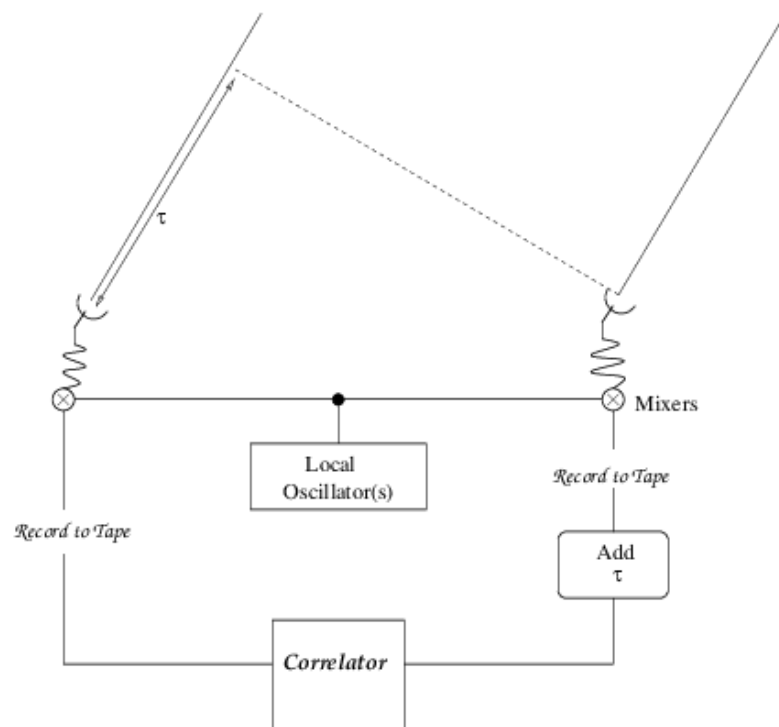
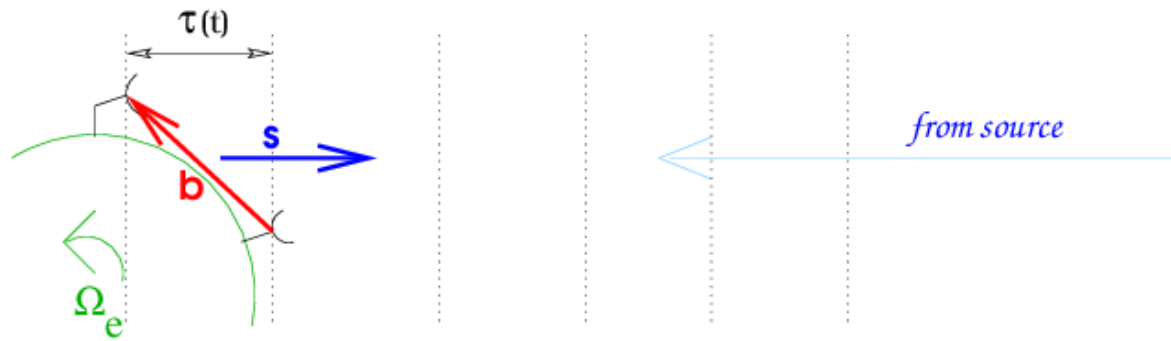


VLBI Correlation 101

Bob Campbell, JIVE
(Joint Institute for VLBI ERIC)

- Concepts via XF
 - Delay tracking
 - Fringe rotation
 - Bsln → Sta Based
- The shift to FX
 - SFXC at JIVE





$$\tau(t) = -\frac{1}{c} \mathbf{B}(t) \cdot \hat{\mathbf{s}}$$

$$\varphi(t) = \nu \tau(t)$$

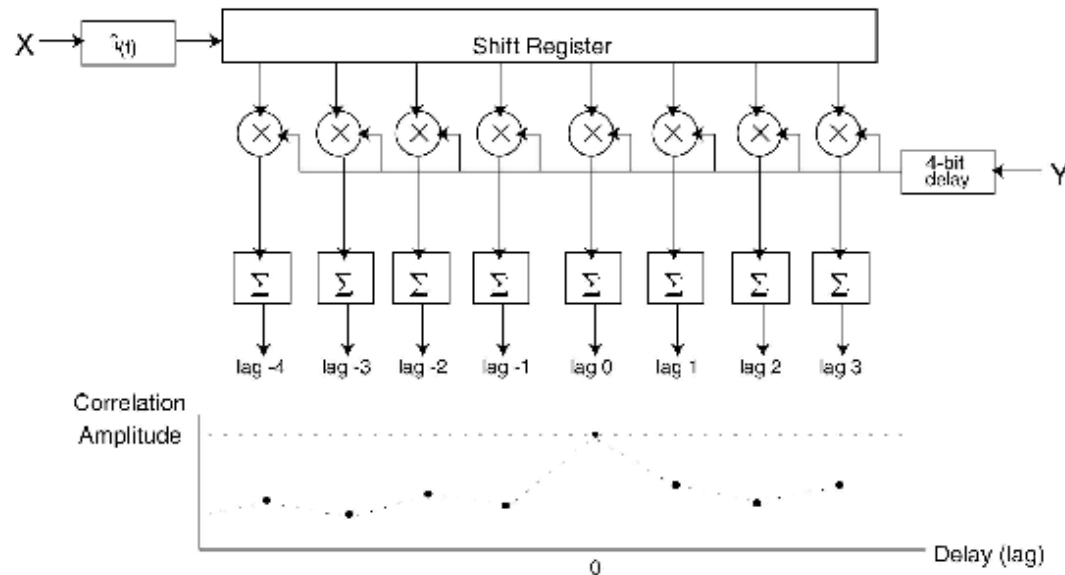
$$\tau_{\max} = 21.3 \text{ ms} \quad (106 \text{ M}\lambda)_{\text{C.band}}$$

$$\dot{\tau}_{\max} = 1.55 \text{ } \mu\text{s/s} \quad (7700 \text{ cyc/s})_{\text{C.band}}$$

Model Constituents (& approx. scales)

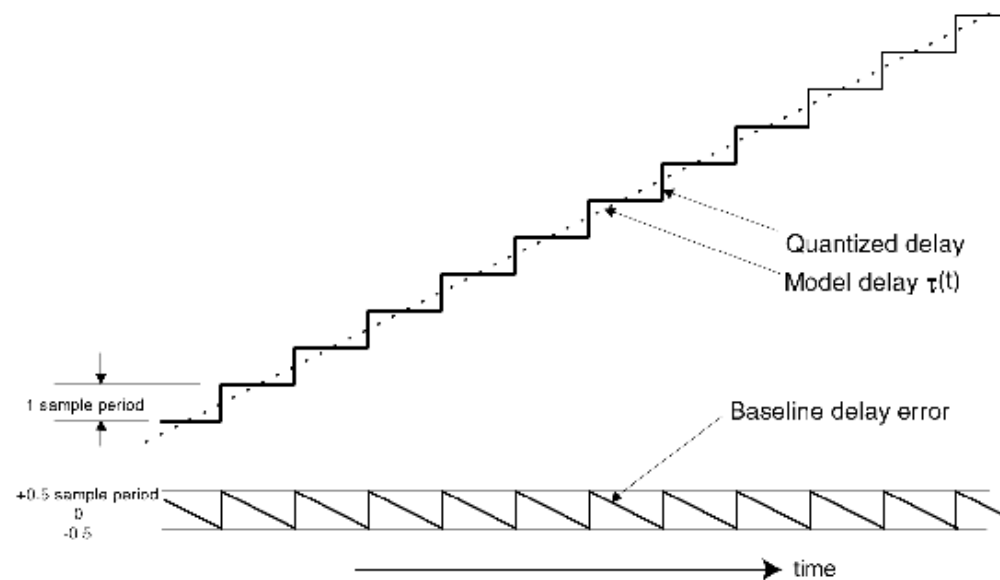
- Station / Source positions: different frames (ITRF, ICRF), motions
- Times: UTC; TAI, TT; UT1; TDB/TCB/TCG
- Orientation: Precession ($50''/\text{yr}$), Nutation ($9.6''$, 18yr), Polar Motion ($0.6''$, 1yr)
- Diurnal Spin: Oceanic friction (2ms/cy), CMB (5ms, dc/ds), AAM (2ms, yrs)
- Tides: Solid-earth (30cm), Pole (2cm)
- Loading: Ocean (2cm), Hydrologic (8mm), Atmospheric (2cm), PGR (mm's/yr)
- Antennas: Axis offset, Tilt, Thermal expansion
- Propagation: Troposphere (dry [7ns], wet [0.3ns]), Ionosphere
- Relativistic $\tau(t)$ calculation: Gravitational delay, Frame choice/consistency

Simplified Real (baseline) Correlator



- Shifting signals from 2 stations by a series of 1-sample steps ($\equiv$ LAGS)
 - Each lag = 1 sampling interval = $\frac{1}{2BW}$ (Nyquist)
- One signal has been delayed by τ (from *a priori* model)
- This delay is constrained to be integral lags (drop/add a sample)

(baseline) Delay Tracking

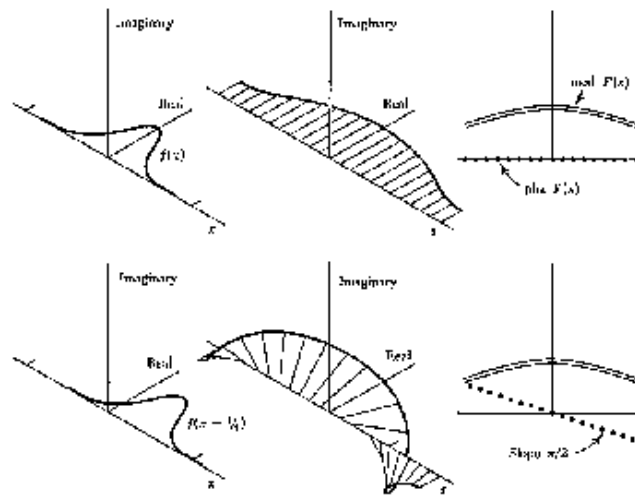


- Delay $\sim$ diurnal sinusoid
 - Amp = B/c ; ~ 3.3 ms per 1000km of baseline
 - Delay steps = $\Delta_\tau = \frac{1}{2BW}$ (1 lag, Nyquist sampled)
- Sawtooth(y) delay error $|\delta_\tau| \leq 0.5$ lag $(\leq \frac{1}{4BW})$
- Adding/subtracting delay steps handled in bitstream

Delay error $\leftrightarrow$ ϕ -slope across band

★ Error = 1 lag $\rightarrow$ 180° edge-to-edge phase slope across band

- Shift theorem: $f(t - \Delta) \Rightarrow e^{-i2\pi(\frac{\nu}{N_{\text{lag}})}\Delta} F(\nu) = e^{-i\pi(\frac{\nu}{N_{\nu}})\Delta} F(\nu)$

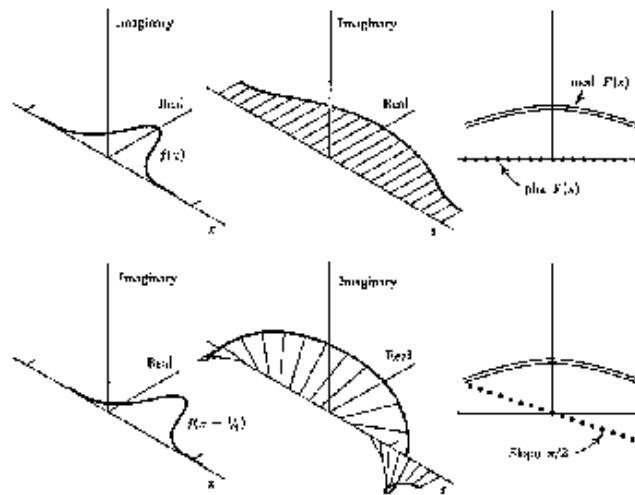


- Can also derive via:
 - Tracing expression of fringe phase (TMS §9.3.i)
 - group-delay = $\frac{1}{c} \int \mu_g ds = \frac{1}{c} \int \frac{d}{d\omega} (\omega \mu_p) ds$

Delay error $\leftrightarrow$ ϕ -slope across band

★ Error = 1 lag $\rightarrow$ 180° edge-to-edge phase slope across band

- Shift theorem: $f(t - \Delta) \Rightarrow e^{-i2\pi(\frac{\nu}{N_{\text{lag}})}\Delta} F(\nu) = e^{-i\pi(\frac{\nu}{N_{\nu}})\Delta} F(\nu)$

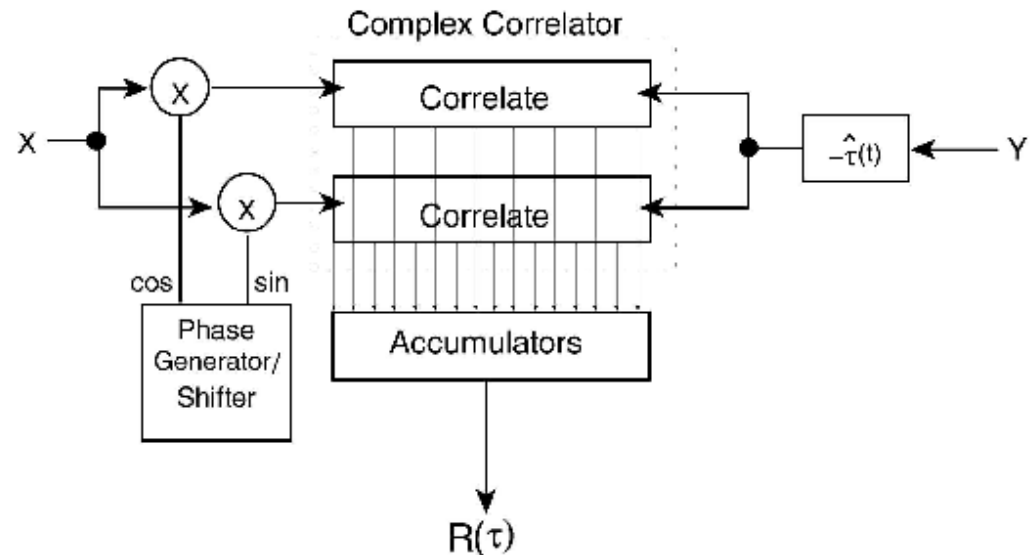


- Can also derive via:
 - Tracing expression of fringe phase (TMS §9.3.i)
 - group-delay = $\frac{1}{c} \int \mu_g ds = \frac{1}{c} \int \frac{d}{d\omega} (\omega \mu_p) ds$

Baseline Complex Correlator

- Delay τ applied to one station
- Fringe rotation (rate $\dot{\tau}$) applied to the other
- cos/sin mixing of fringe rotation (Hilbert transforms: Bracewell Ch.12)
- output = one complex number per lag

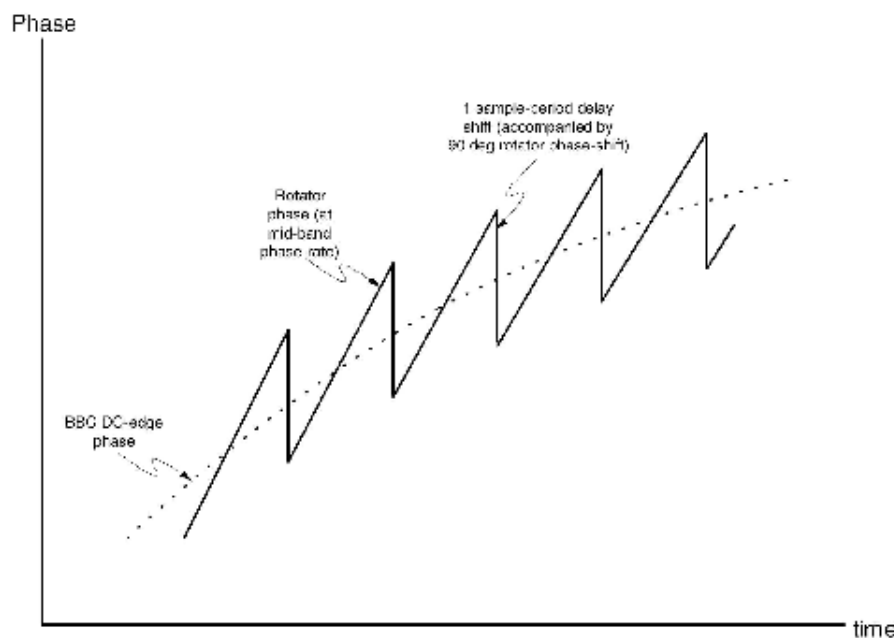
- **XF**



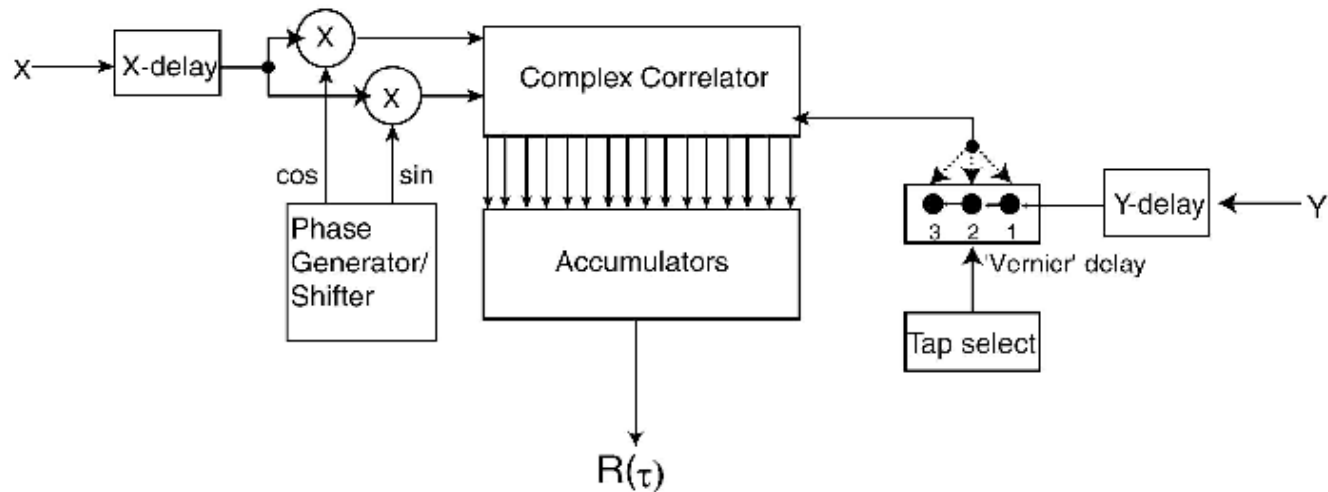
Running the fringe rotator

- track φ at band center (minimize phase offsets at edges)
- DC edge $\rightarrow$ band center: φ_{FR} increment $= \pm 2\pi(\frac{BW}{2})\delta_\tau = \pm \pi BW \delta_\tau$
- Step fringe rotation $\pm 90^\circ$ at each instant delay shifted by $\pm \Delta_\tau$

$$\delta_\tau \rightarrow \delta_\tau \pm \Delta_\tau = \delta_\tau \pm \frac{1}{2BW}, \quad \varphi_{FR} \rightarrow \varphi_{FR} \pm \frac{\pi}{2}$$

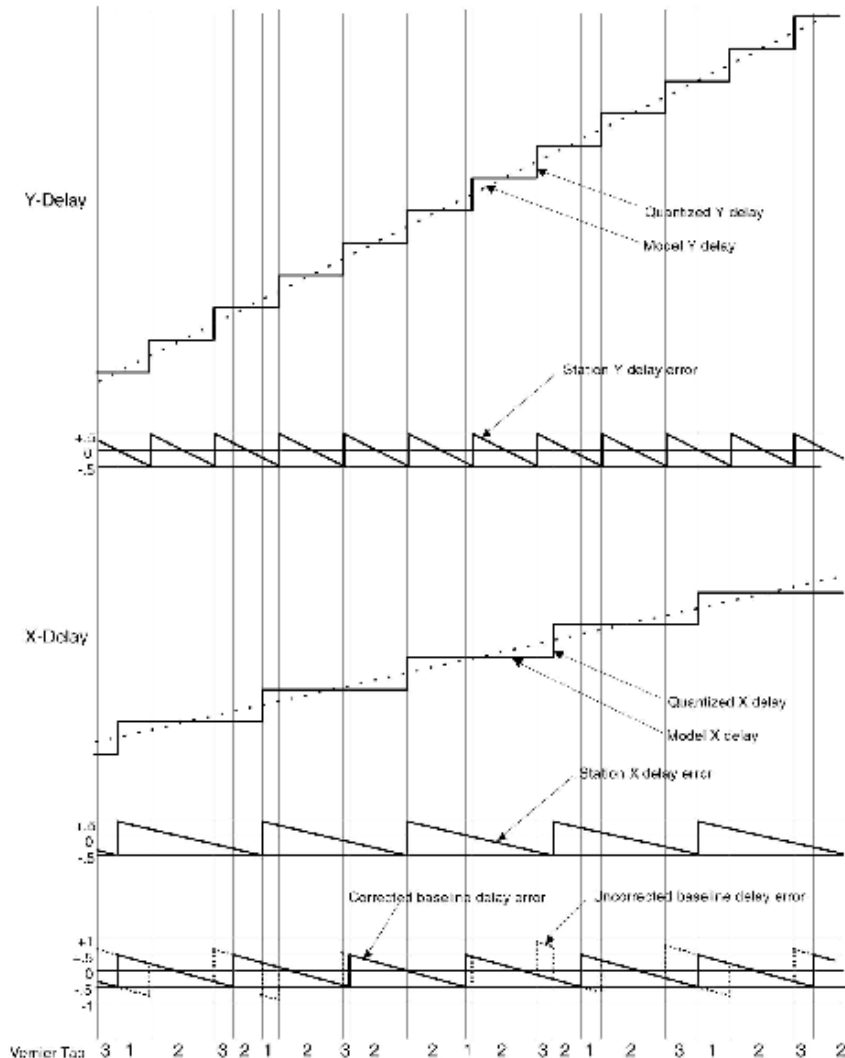


Station-based Correlator



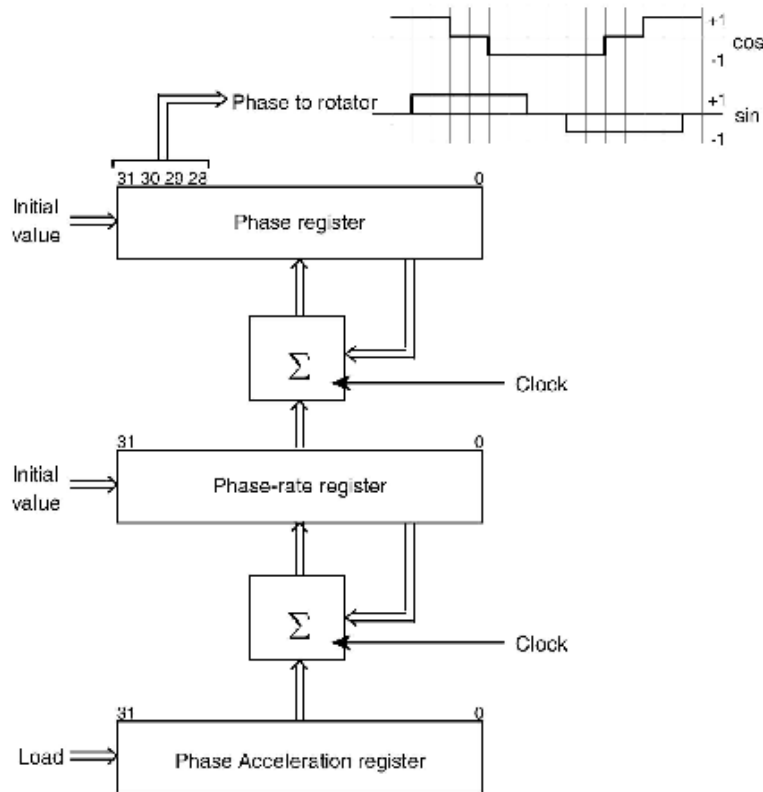
- Same as baseline XF, **PLUS** the “Vernier” delay tap

Station-based Delay Tracking



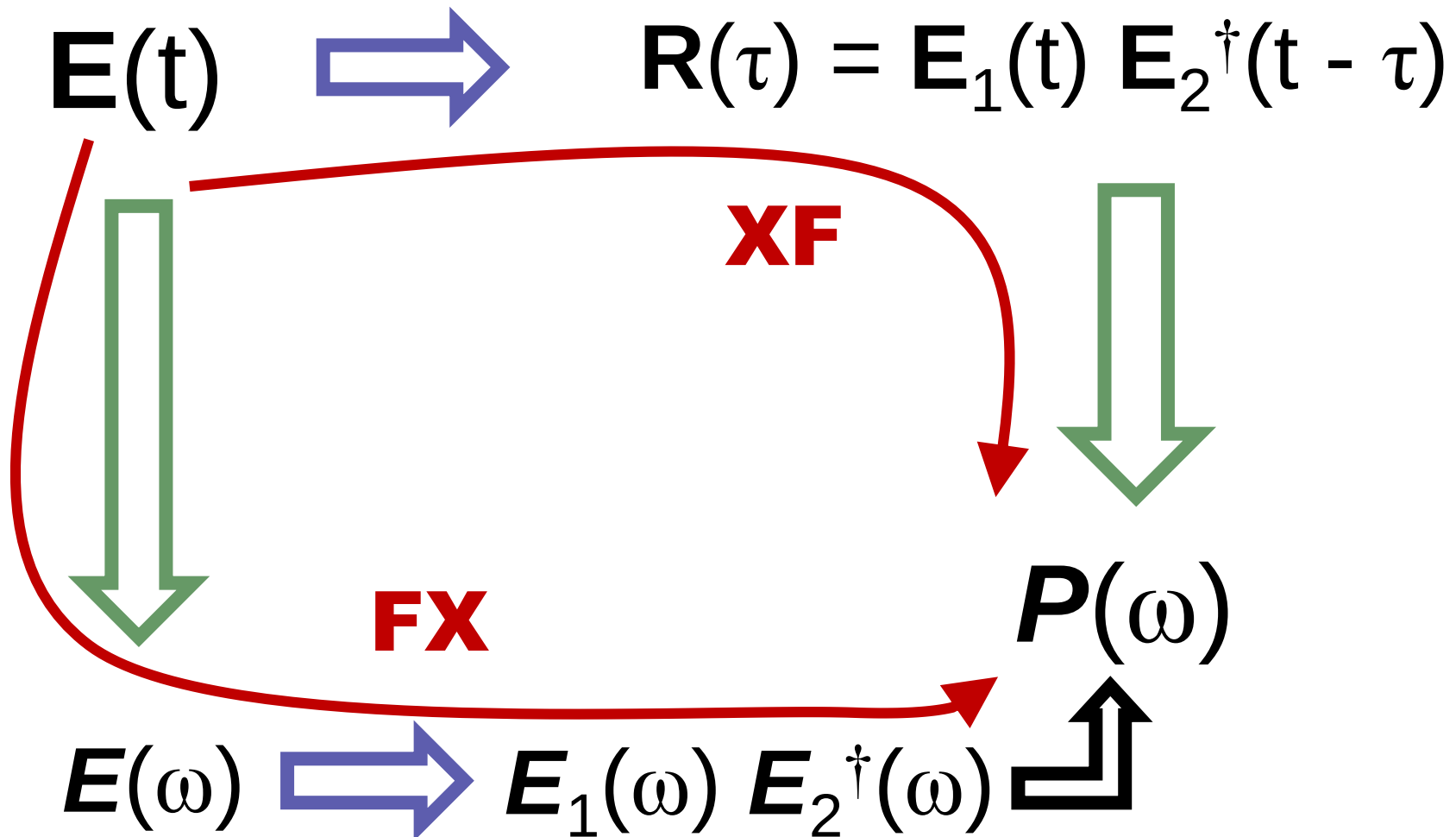
- $\exists$ periods of $|\delta\tau| > 0.5$ lag
 - Worse phase slopes accross band
- Vernier delay tap (± 1 lag) set to shift $|\delta\tau| > 0.5$ lag to $|\delta\tau| \leq 0.5$ lag

XF fringe-rotate both stations?

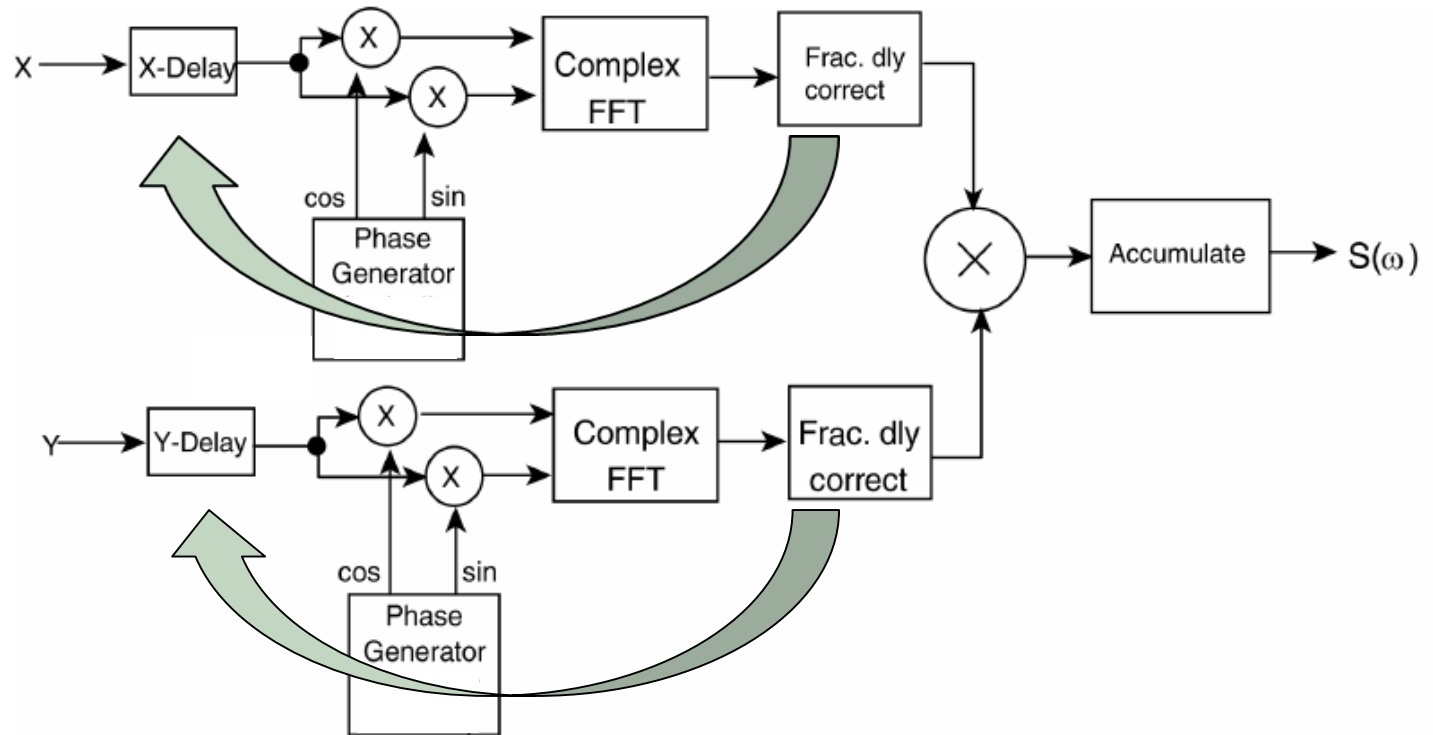


- 3-level sinusoids (+1, 0, -1)
 - doubled SNR loss
- Energy scattered into higher harmonics
- Fringe-rotating both stations would create two sets of spurious harmonics that correlate with each other
 - Delay-rate ~ 0 events
- Updating phase “integration”
 - Speed-up
 - Oversampling

Correlations & Fourier Transforms



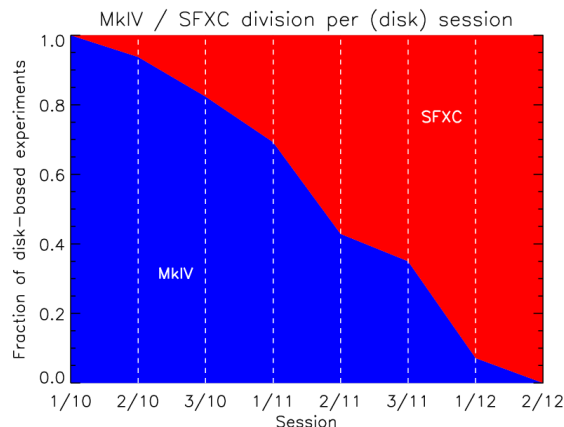
FX Station-based Correlator (SFXC)



- Fractional bit-shift delay via separate FFT earlier in delay compensation
- Quadrature rotation for removing *a priori* fringe rate
- A windowing step performed between fringe rate and station-based FFT
- The size of the 2 FFT's independent of each other.

Correlator Evolution at JIVE (I)

- ENV MkIV Data Processor at JIVE
 - 16 stations (via tape, later Mark5 disk-pack playback)
 - Max. 1Gbps (but runs real-time regardless of input rate <1Gbps)
 - Max. 2048 frequency points per subband/pol
 - Byzantine rules for joint max. $N\{\text{subbands, pols, frequency points}\}$
 - Bespoke (irreplaceable) chips
- EVN Software Correlator at JIVE (SFXC)
 - Much more flexible, but not necessarily real-time
 - 1st user experiment correlated on SFXC from May/Jun 2010



Correlator Evolution at JIVE (II)

- Astronomy Gains enabled by SFXC
 - ~Arbitrary number of stations, total bit-rate, channel BW
 - Minimum $t_{\text{int}} = 1$ correlation FFT size in lag-space
 - Wide-field mapping (large $N_{\text{freq.pt.}}$ per SB/pol, short t_{int})
 - Multiple output phase centers per pointing
 - Delay/rate steering to N phase centers from within “internal” wide-field correlation
 - Pulsar gating/binning (1 gate, N equally spaced bins + out-of-gate bin)
 - Incoherent and coherent de-dispersion
 - Spectral zooming
 - Choice of spectral-windowing function (Hann, Hamming, top-hat, cosine)
 - Mixed input channel BWs → handling heterogeneous back-ends
 - External model files (e.g., near-field target or orbiting antenna)
 - Pure station-based fringe rotation to CoE; consistent cross-pols

Useful Links

- Correlators in general (slants towards geodetic perspective)
 - Whitney 2000, *IVS General Meeting 2000*

- EVN Correlators
 - MarkIV (XF)
 - Schilizzi et al. 2001, *Exper. Astron.*, 12, 49 (EVN MarkIV)
 - Whitney et al. 2004, *Rad. Sci.*, 39, 1007 (Archit. & Algo.'s)
 - SFXC
 - Keimpema et al. 2015, *Exper. Astron.*, 39, 259
 - Keimpema et al. 2025, *Exper. Astron.*, in press